Dynamic stability of the new planetary system HD 73526, in 2:1 mean motion resonance, is not without controversy. Whereas this system has been found at first stable for coplanar orbits over a 1 Myr test integration, techniques of global dynamics analysis show a quite unstable behavior within the strict observational data. Nevertheless, we find a stabilizing mechanism involving the undetermined observational parameters, namely the orbital inclinations and longitudes of nodes. Phase portrait of the system HD 73526 suggests that after its formation, it could have pass a phase of dynamical evolution driving it towards the present equilibrium.